

Numerical Simulation of Anisotropic Mean Curvature of Graphs in Relative Geometry

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Abstract: The aim of this paper is the numerical simulation of anisotropic mean curvature of graphs in the context of relative geometry, developed in [1]. We extend results in [4] to our problem; we prove an existence theorem and energy equality. The numerical scheme is based on the method of lines where the spatial derivatives are approximated by finite differences [2]. We then solve the resulting ODE system by means of the adaptive Runge-Kutta-Merson method. To show the stability of the scheme we prove the discrete version of the energy equality. Finally, we show experimental order of convergence and results of numerical experiments with various anisotropy settings.

Keywords: anisotropy; mean curvature; Finsler geometry; method of lines; FDM

1 Introduction

The paper studies the following motion law for surfaces in $\mathbb{R}^3$ denoted by Γ :

$$\text{velocity} = \text{curvature} + \text{forcing} \quad (1)$$

in a certain sense which is specified below. Both the velocity and the curvature are evaluated with respect to the direction given by a vector locally influenced by the orientation of the Euclidean normal vector to Γ .

One example of the law (1) is represented by the isotropic mean-curvature flow given by the equation

$$v_\Gamma = -\kappa_\Gamma + f \text{ on } \Gamma(t) \quad (2)$$

in the direction of n_Γ which is the Euclidean normal vector to Γ , while v_Γ the normal velocity, κ_Γ the mean curvature, and f the forcing term. The equation (2) in the form of the Gibbs-Thompson law is contained in the modified Stefan problem. For details, we refer the reader to [9, 16].

One of few anisotropic examples where the analytical solution is known considers a ball under the relative geometry which shrinks according to (1) with $f = 0$. In this case we have the initial ball with radius r_0 , normal velocity $\dot{r}$, actual curvature along the ball of radius r being $\frac{1}{r}$. The equation (1) reads

$$\dot{r} = -\frac{1}{r},$$

and has the solution

$$r(t) = \sqrt{r_0^2 - 2t}.$$

This law has been intensively studied, see e.g. [4, 5, 13].

This paper deals with the motion by anisotropic mean curvature in relative geometry associated with the Finsler metric, developed in [1], which reads

$$v_{\Gamma, \phi} = -\kappa_{\Gamma, \phi} + f \text{ on } \Gamma(t) \quad (3)$$

Here, $v_{\Gamma, \phi}$ denotes the normal velocity, $\kappa_{\Gamma, \phi}$ is the anisotropic mean curvature of $\Gamma(t)$ with respect to the Finsler metric ϕ , and f is the forcing term.

Deckelnick and Dziuk proved the convergence and gave the optimal error estimates using finite element method for graph [4, 7] and parametric case [8]. Haußer and Voigt [11] presented a parametric finite element approximation for a regularized version. Pozzi studied the anisotropic mean curvature flow in higher codimension in [15].

2 Anisotropy in Relative Geometry

In what follows we shall first define anisotropy by means of the Finsler geometry; then, we shall transform the motion law (3) into graph formulation. For this purpose, we assume that there is a smooth function with non-vanishing gradient $p: \mathbb{R}^{2+1} \rightarrow \mathbb{R}$ such that

$$\Gamma(t) = \{[x, y] \in \mathbb{R}^3 \mid y = p(t, x), x \in \Omega \subset \mathbb{R}^2\}.$$

We say that a continuous function $\phi: \mathbb{R}^3 \rightarrow \mathbb{R}_0^+$ is a Finsler metric if it satisfies the properties

1. $\phi \in C^{3+\alpha}(\mathbb{R} - \{0\})$,
 2. ϕ^2 is strictly convex,
 3. $\phi(t\eta) = |t|\phi(\eta)$, $t \in \mathbb{R}$, $\eta \in \mathbb{R}^3$,
 4. $\lambda|\eta| \leq \phi(\eta) \leq \Lambda|\eta|$, $\eta \in \mathbb{R}^3$,
- for two suitable constants $0 < \lambda \leq \Lambda < \infty$.

Associated to ϕ we define the unit ball (also so-called Wulff shape)

$$B_\phi = \{\eta \in \mathbb{R}^3 | \phi(\eta) \leq 1\}.$$

One can prove that a dual function $\phi^0: \mathbb{R}^3 \rightarrow \mathbb{R}_0^+$ given by

$$\phi^0(\eta^*) = \sup\{\eta^* \cdot \eta | \eta \in B_\phi\}$$

is also a Finsler metric.

For simplicity we use η instead of η^* . Then the following relations hold [3]

$$\phi_\eta^0(t\eta) = \frac{t}{|t|} \phi_\eta^0(\eta), \quad \phi_{\eta\eta}^0(t\eta) = \frac{1}{|t|} \phi_{\eta\eta}^0(\eta), \quad t \in \mathbb{R} - \{0\},$$

$$\zeta \cdot \phi^0(\eta) \phi_{\eta\eta}^0(\eta) \cdot \zeta \geq \gamma_0 \left| \zeta - \frac{\zeta \cdot \eta}{|\eta|^2} \eta \right|^2, \quad \eta \neq 0, \quad \zeta \in \mathbb{R}^3, \quad \gamma_0 > 0,$$

where the index η means the derivative with respect to η .

We define the map $T^0: \mathbb{R}^3 \rightarrow \mathbb{R}^3$ as

$$T^0(\eta) = (\tilde{T}^0(\eta), T_3^0(\eta)) = \phi^0(\eta) \phi_\eta^0(\eta), \quad \eta \neq 0,$$

$$T^0(0) = 0.$$

Then, the ϕ -normal vector, ϕ -mean curvature, and ϕ -normal velocity of Γ are defined as

$$n_{\Gamma, \phi} = \frac{T^0(\nabla p, -1)}{\phi^0(\nabla p, -1)} = \phi_\eta^0(\nabla p, -1), \quad (4)$$

$$\kappa_{\Gamma, \phi} = \operatorname{div} n_{\Gamma, \phi} = \nabla \cdot \frac{\tilde{T}^0(\nabla p, -1)}{\phi^0(\nabla p, -1)}, \quad (5)$$

$$v_{\Gamma, \phi} = -\frac{\partial_t p}{\phi^0(\nabla p, -1)}. \quad (6)$$

By substituting the quantities (4)-(6) into the Eq. (3), we obtain the non-linear parabolic partial differential equation

$$\partial_t p = \phi^0(\nabla p, -1) \left(\nabla \cdot \left(\frac{\tilde{T}^0(\nabla p, -1)}{\phi^0(\nabla p, -1)} \right) + f \right) \text{ on } \Omega \times (0, T). \quad (7)$$

The initial and boundary conditions are given by

$$p|_{t=0} = p_0 \text{ on } \bar{\Omega}, \quad (8)$$

$$p = p_0 \text{ on } \partial\Omega \times (0, T). \quad (9)$$

In our numerical experiments we use the Finsler metrics listed below. We denote $n = \frac{\eta}{|\eta|}$. The corresponding Wulff shapes are illustrated in Fig. 1.

The 4-fold anisotropy reads as

$$\phi(\eta) = |\eta| \left(1 - A_1 (1 - (n_1^4 + n_2^4 + n_3^4)) \right). \quad (10)$$

The 6-fold anisotropy reads as

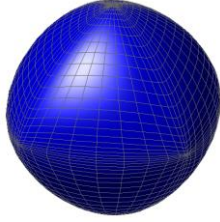
$$\begin{aligned} \phi_\eta^0(\eta) = |\eta| & \left(1 - A_1 \left(n_1^4 + n_2^4 + n_3^4 - \frac{3}{5} \right) \right. \\ & \left. + A_2 \left(3(n_1^4 + n_2^4 + n_3^4) + 66n_1^2 n_2^2 n_3^2 - \frac{17}{7} \right) \right) \end{aligned} \quad (11)$$

The 8-fold anisotropy reads as

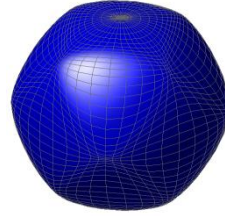
$$\begin{aligned} \phi_\eta^0(\eta) = |\eta| & (1 - A_1 (n_1^8 + n_2^8 + n_3^8 \\ & - 28(n_1^6 n_2^2 + n_1^2 n_2^6 + n_2^6 n_3^2 + n_2^2 n_3^6 + n_3^6 n_1^2 + n_3^2 n_1^6) \\ & + 70(n_1^4 n_2^4 + n_2^4 n_3^4 + n_3^4 n_1^4))). \end{aligned} \quad (12)$$

The regularized l_1 -anisotropy reads as

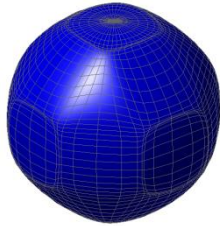
$$\phi_\eta^0(\eta) = \sum_{i=1}^3 \left(\eta_i^2 + A_1 \sum_{j=1}^3 \eta_j^2 \right)^{\frac{1}{2}}. \quad (13)$$



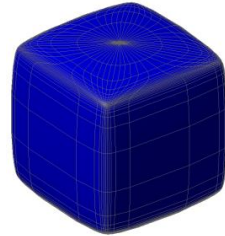
4-fold anisotropy, $A_1 = 0.24$



6-fold anisotropy, $A_1 = -0.035$, $A_2 = -0.035$



8-fold anisotropy, $A_1 = -0.015$



regularized l_1 -norm, $A_1 = 0.01$

Figure 1
Wulff shapes for various anisotropies

3 Analytical Properties

In the following section we shall introduce some analytical results for law (7) in the context of relative geometry, which are due to [4, 11, 12]. We shall prove the energy equality and give the existence result for our problem.

Theorem 1. *For the solution of problem (7)-(9), one has the energy equality*

$$\int_{\Omega} p_t \left(\frac{p_t}{\phi^0(\nabla p, -1)} - f \right) + \frac{d}{dt} \int_{\Omega} \phi^0(\nabla p, -1) = 0.$$

If $f = 0$, then

$$\int_{\Omega} \frac{p_t^2}{\phi^0(\nabla p, -1)} + \frac{d}{dt} \int_{\Omega} \phi^0(\nabla p, -1) = 0. \quad (14)$$

Proof. Since $p_t = 0$ on $\partial\Omega$, the proof is straightforward

$$\begin{aligned} \frac{d}{dt} \int_{\Omega} \phi^0(\nabla p, -1) &= \int_{\Omega} \phi_{\eta}^0(\nabla p, -1) \cdot [\nabla p, -1]_t = \int_{\Omega} [\nabla p_t, 0] \cdot \frac{T^0(\nabla p, -1)}{\phi^0(\nabla p, -1)} \\ &= \int_{\Omega} \nabla p_t \cdot \frac{\tilde{T}^0(\nabla p, -1)}{\phi^0(\nabla p, -1)} = \int_{\Omega} p_t \nabla \cdot \frac{\tilde{T}^0(\nabla p, -1)}{\phi^0(\nabla p, -1)} \\ &= - \int_{\Omega} p_t \left(\frac{p_t}{\phi^0(\nabla p, -1)} - f \right). \end{aligned}$$

If $f = 0$, we obtain the equality (14).

Lemma 1. *Let $v(\tilde{\eta}) = \phi^0(\tilde{\eta}, -1)$, $a_i = \frac{T_i^0(\nabla p, -1)}{\phi^0(\nabla p, -1)}$, and $a_{ij} = \frac{\partial a_i}{\partial \eta_j}$. Then for the solution of the problem (7)-(9) with $f = 0$ one has the identity*

$$\begin{aligned} v_t &= \sum_{i,j=1}^2 \frac{\partial}{\partial x_i} \left(a_{ij} \frac{\partial v}{\partial x_j} \right) v + \sum_{l=1}^2 \kappa_{\phi} a_l \frac{\partial v}{\partial x_l} \\ &\quad + \sum_{i,j,k=1}^2 a_{ij} a_{kl} \frac{\partial^2 p}{\partial x_i \partial x_k} \frac{\partial^2 p}{\partial x_j \partial x_l} v. \end{aligned} \quad (15)$$

Proof. We have

$$v_t = \phi_{\eta}^0(\nabla p, -1) \nabla p_t = \sum_{l=1}^2 a_l \frac{\partial (v \kappa_{\phi})}{\partial x_l} = v \sum_{l=1}^2 a_l \frac{\partial \kappa_{\phi}}{\partial x_l} + \kappa_{\phi} \sum_{l=1}^2 a_l \frac{\partial v}{\partial x_l}.$$

Let now compute

$$\begin{aligned} \sum_{l=1}^2 a_l \frac{\partial \kappa_\phi}{\partial x_l} &= \sum_{l=1}^2 a_l \frac{\partial}{\partial x_l} \sum_{i=1}^2 \frac{\partial}{\partial x_i} a_i = \sum_{i,l=1}^2 a_l \frac{\partial}{\partial x_i} \frac{\partial}{\partial x_l} a_i \\ &= \sum_{i,l=1}^2 \frac{\partial}{\partial x_i} \left(a_l \frac{\partial}{\partial x_l} a_i \right) - \sum_{i,l=1}^2 \frac{\partial}{\partial x_i} a_l \frac{\partial}{\partial x_l} a_i. \end{aligned}$$

Since

$$\frac{\partial}{\partial x_j} v = \sum_{k=1}^2 a_k \frac{\partial}{\partial x_j} \frac{\partial}{\partial x_k} p, \quad \frac{\partial}{\partial x_l} a_i = \sum_{j=1}^2 a_{ij} \frac{\partial}{\partial x_j} \frac{\partial}{\partial x_l} p,$$

we get the identity (15).

Theorem 2. Let $\partial\Omega \in C^{3+\alpha}$ and $d(x) := \text{dist}(x, \partial\Omega)$. We assume that

$$\sum_{i,j=1}^2 \phi_{\eta_i \eta_j}^0(\nabla d(x), 0) d_{x_i x_j}(x) \leq 0, \quad x \in \partial\Omega.$$

Let $p_0 \in C^{3+\alpha}(\bar{\Omega})$ satisfies the compatibility condition

$$\sum_{i,j=1}^2 \phi_{\eta_i \eta_j}^0(\nabla p_0, -1) p_{0, x_i x_j} = 0, \quad x \in \partial\Omega.$$

Then (7)-(9) with $f = 0$ has a solution $p \in H^{3+\alpha, \frac{3+\alpha}{2}}(\bar{\Omega} \times [0, T])$ with $p_t \in L^2(0, T; H^{2,2}(\Omega))$ for all $T < \infty$.

Proof. Similarly as in [4] we are looking for a solution of the initial boundary value problem

$$p_t - \sum_{i,j=1}^2 a_{ij}(\nabla p) p_{x_i} p_{x_j} = 0$$

but with the difference

$$a_{ij}(\tilde{\eta}) = \phi^0(\tilde{\eta}, -1) \phi_{\eta_i \eta_j}^0(\tilde{\eta}, -1).$$

Since ϕ^0 is a Finsler metric, $\phi^0 \in C^{0,1}(\mathbb{R})$ holds. Moreover, since $\phi^0 \in C^{3+\alpha}(\mathbb{R} - \{0\})$, we have $a_{ij} \in C^{0,1}(\mathbb{R})$.

Following standard lines of Theorem 4.1 in [4] and using the previous Lemma 1 we can show there is a constant K such that for every solution p^τ of

$$p_t^\tau - \tau \sum_{i,j=1}^2 a_{ij}(\nabla p^\tau) p_{x_i x_j}^\tau - (1 - \tau) \Delta p^\tau = 0 \quad \text{in } \Omega \times (0, T),$$

$$p^\tau = \tau p_0 \quad \text{on } \partial\Omega \times (0, T),$$

$$p^\tau(\cdot, 0) = \tau p_0 \quad \text{on } \Omega,$$

the estimate

$$\max_{\Omega \times (0, T)} |p^\tau| + \max_{\Omega \times (0, T)} |\nabla p^\tau| \leq K$$

is valid. This means (7)-(9) with $f = 0$ has a solution $p \in H^{3+\alpha, \frac{3+\alpha}{2}}(\bar{\Omega} \times [0, T])$.

4 Numerical Scheme

We employed the numerical scheme based on the method of lines. The spatial derivatives are discretized and the time variable is left continuous. After discretizing the problem by finite differences in space, we solve the resulting ODE system by the adaptive Runge-Kutta-Merson method. We consider the computational domain $\Omega = (0, L_1) \times (0, L_2)$ and introduce the following notation:

$$h_1 = \frac{L_1}{N_1}, h_2 = \frac{L_2}{N_2}$$

$$\omega_h = \{[ih_1, jh_2] | i = 1, \dots, N_1 - 1; j = 1, \dots, N_2 - 1\},$$

$$\bar{\omega}_h = \{[ih_1, jh_2] | i = 0, \dots, N_1; j = 0, \dots, N_2\},$$

$$\gamma_h = \omega_h - \bar{\omega}_h,$$

$$u_{ij} = u(ih_1, jh_2),$$

$$u_{x_1, ij} = \frac{u_{i+1, j} - u_{ij}}{h_1}, u_{x_2, ij} = \frac{u_{i, j+1} - u_{ij}}{h_2},$$

$$u_{\bar{x}_1, ij} = \frac{u_{ij} - u_{i-1, j}}{h_1}, u_{\bar{x}_2, ij} = \frac{u_{ij} - u_{i, j-1}}{h_2},$$

$$\nabla_h u = [u_{x_1}, u_{x_2}],$$

$$\bar{\nabla}_h u = [u_{\bar{x}_1}, u_{\bar{x}_2}],$$

$$\mathcal{P}_h g = g|_{\bar{\omega}_h}.$$

We define the following expressions

$$\begin{aligned}
 (f, g)_h &= \sum_{i=1}^{N_1-1} \sum_{j=1}^{N_2-1} h_1 h_2 f_{ij} g_{ij}, \quad \|f\|_h^2 = (f, f)_h, \\
 (f^1, g^1] &= \sum_{i=1}^{N_1-1} \sum_{j=1}^{N_2-1} h_1 h_2 f_{ij}^1 g_{ij}^1, \\
 (f^2, g^2] &= \sum_{i=1}^{N_1-1} \sum_{j=1}^{N_2-1} h_1 h_2 f_{ij}^2 g_{ij}^2, \\
 (\mathbf{f}, \mathbf{g}] &= (f^1, g^1] + (f^2, g^2], \\
 (f, g] &= \sum_{i=1}^{N_1} \sum_{j=1}^{N_2} h_1 h_2 f_{ij} g_{ij}.
 \end{aligned}$$

We then propose a semi-discrete scheme [2]

$$\begin{aligned}
 p_t^h &= \phi^0(\bar{\nabla}_h p^h, -1) \left(\nabla_h \cdot \left(\frac{\tilde{T}^0(\bar{\nabla}_h p^h, -1)}{\phi^0(\bar{\nabla}_h p^h, -1)} \right) + f \right) \text{on } \omega_h \times (0, T), \\
 p^h|_{t=0} &= \mathcal{P}_h p_0 \text{on } \bar{\omega}_h \\
 p^h &= \mathcal{P}_h p_0 \text{on } \gamma_h \times (0, T).
 \end{aligned} \tag{16}$$

This is an ODE system and existence and uniqueness of solutions are guaranteed by the theory of ordinary differential equations (the Picard–Lindelöf theorem).

As the stability criterion we use the basic energy equality (14). For this purpose we shall now prove the discrete version of Theorem 1.

Theorem 3. For the solution of problem (16), the following energy equality holds

$$\left(p_t^h, \frac{p_t^h}{\phi^0(\bar{\nabla}_h p^h, -1)} - f \right)_h + \frac{d}{dt} (\phi^0(\bar{\nabla}_h p^h, -1), 1] = 0.$$

If $f = 0$, then

$$\left((p_t^h)^2, \frac{1}{\phi^0(\bar{\nabla}_h p^h, -1)} \right)_h + \frac{d}{dt} (\phi^0(\bar{\nabla}_h p^h, -1), 1] = 0. \tag{17}$$

Proof. Applying the grid version of Green's formula as in [2], we obtain

$$\begin{aligned}
 \left(p_t^h, \frac{p_t^h}{\phi^0(\bar{\nabla}_h p^h, -1)} - f \right)_h &= \left(p_t^h, \nabla_h \cdot \left(\frac{\tilde{T}^0(\bar{\nabla}_h p^h, -1)}{\phi^0(\bar{\nabla}_h p^h, -1)} \right) \right)_h \\
 &= \left(\bar{\nabla}_h p_t^h, \frac{\tilde{T}^0(\bar{\nabla}_h p^h, -1)}{\phi^0(\bar{\nabla}_h p^h, -1)} \right] \\
 &= - \sum_{i=1}^{N_1} \sum_{j=1}^{N_2-1} h_1 h_2 p_t^h|_{\bar{x}_{1,ij}} \frac{T_1^0(\bar{\nabla}_h p^h, -1)}{\phi^0(\bar{\nabla}_h p^h, -1)} \Big|_{ij}
 \end{aligned}$$

$$\begin{aligned}
& - \sum_{i=1}^{N_1-1} \sum_{j=1}^{N_2} h_1 h_2 p_t^h|_{\bar{x}_2, ij} \frac{T_2^0(\bar{\nabla}_h p^h, -1)}{\phi^0(\bar{\nabla}_h p^h, -1)} \Big|_{ij} \\
& = - \sum_{i=1}^{N_1} \sum_{j=1}^{N_2} h_1 h_2 p_t^h|_{\bar{x}_1, ij} \frac{T_1^0(\bar{\nabla}_h p^h, -1)}{\phi^0(\bar{\nabla}_h p^h, -1)} \Big|_{ij} \\
& - \sum_{i=1}^{N_1} \sum_{j=1}^{N_2} h_1 h_2 p_t^h|_{\bar{x}_2, ij} \frac{T_2^0(\bar{\nabla}_h p^h, -1)}{\phi^0(\bar{\nabla}_h p^h, -1)} \Big|_{ij} = - \frac{d}{dt} (\phi^0(\bar{\nabla}_h p^h, -1), 1].
\end{aligned}$$

If $f = 0$, we get the equality (17).

5 Computational Results

We first investigate the convergence of the numerical scheme. Then, we explore the long time behaviour of the anisotropic motion law (3).

Experimental order of convergence. The computations have been performed over a range of different grid resolutions which allows quantifying the numerical convergence by the experimental order of convergence (EOC). A numerical solution computed on the finest grid is used to substitute the analytical solution. Given errors Error_1 and Error_2 for two mesh sizes h_1 , h_2 , respectively, the EOC is defined as

$$\text{EOC} = \frac{\log(\text{Error}_1/\text{Error}_2)}{\log(h_1/h_2)}.$$

The result is shown in the following table.

Table 1
Experimental order of convergence of the scheme (16)

N	h	Error L_∞	EOC L_∞	Error L_2	EOC L_2
50	1/50	0.05924	-	0.01175	-
100	1/100	0.03676	0.68843	0.00731	1.00000
150	1/150	0.02689	0.77219	0.00511	1.00000
200	1/200	0.02058	0.92781	0.00357	1.00000

Morphology evolution. We present the solutions at different times for various anisotropies. Figs. 2-6 show surface evolutions under anisotropic mean curvature flow without the forcing term ($f = 0$). Anisotropy is shown to be crucial in the formation of different surface morphologies. The surface is first determined by symmetry of anisotropy; it then evolves towards to the flat surface. Finally, the effect of the forcing term f on the surface evolution is shown in Fig. 7.

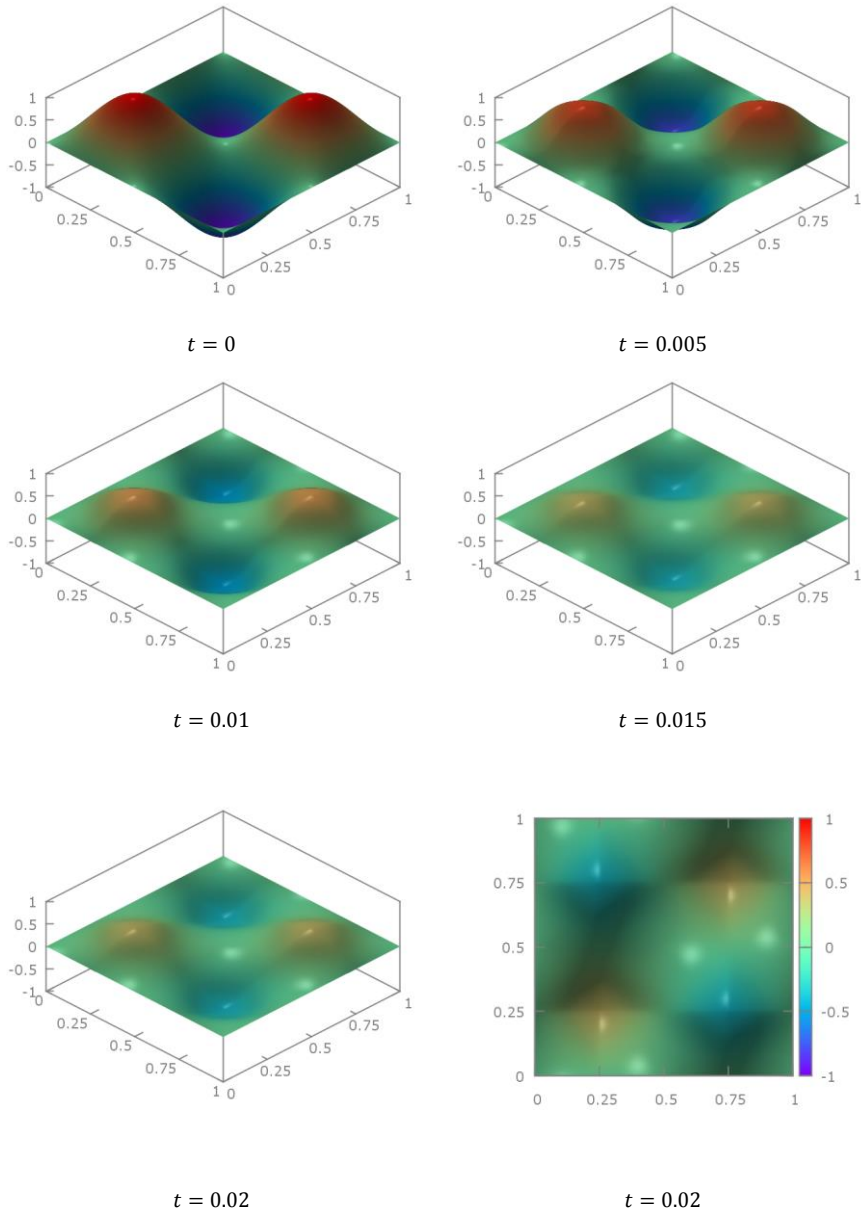


Figure 2
Morphology evolution for $f = 0$, the 4-fold anisotropy (10) with $A_1 = 0.24$,
 $p_0(x_1, x_2) = \sin(2x_1)\sin(2x_2)$ at different times

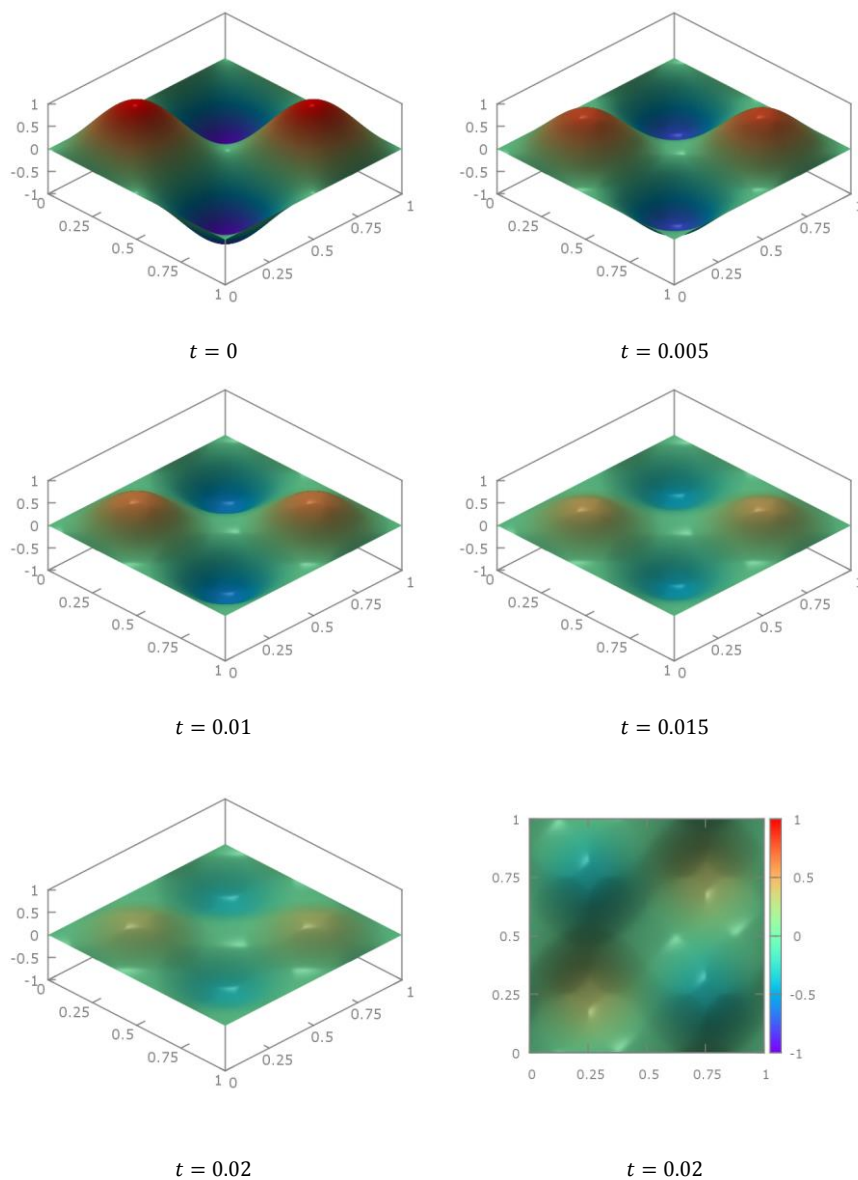


Figure 3

Morphology evolution for $f = 0$, the 6-fold anisotropy (11) with $A_1 = -0.035$, $A_2 = -0.035$,
 $p_0(x_1, x_2) = \sin(2x_1)\sin(2x_2)$ at different times

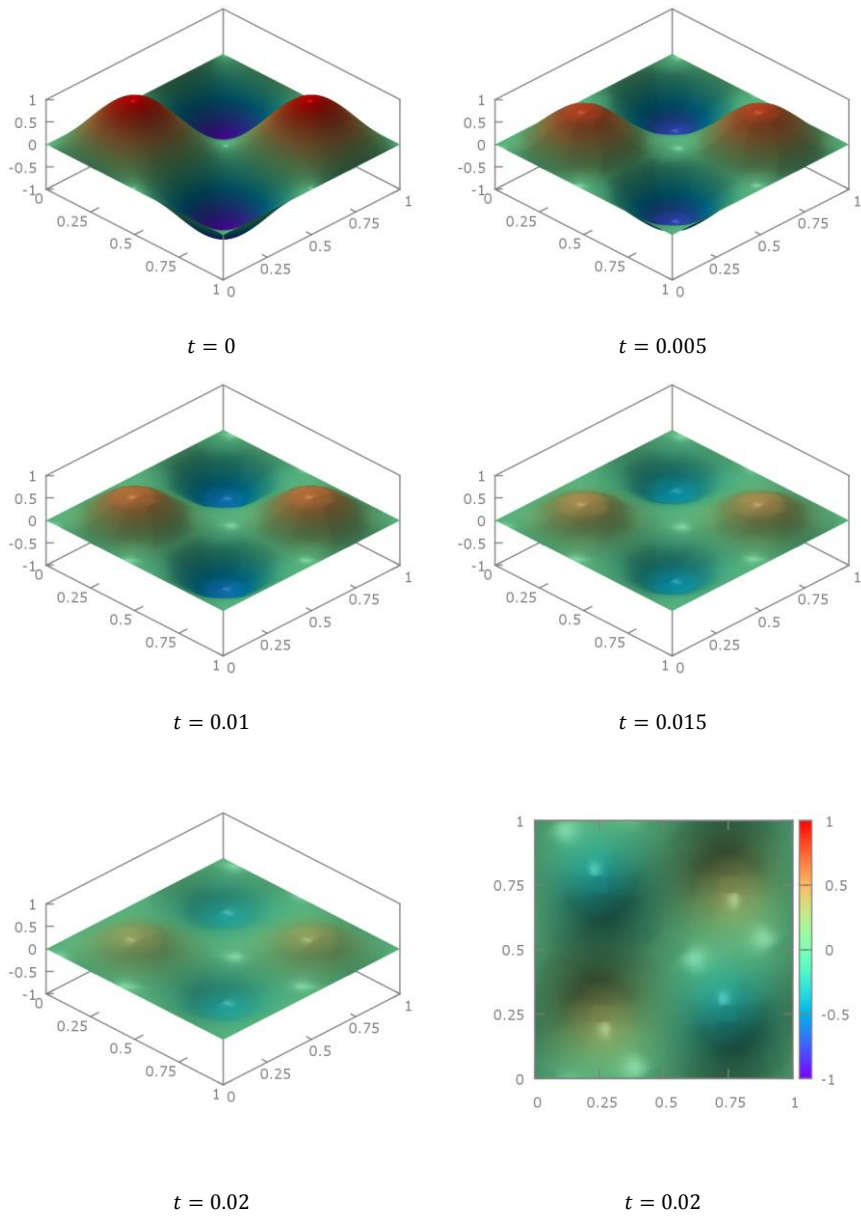


Figure 4
Morphology evolution for $f = 0$, the 8-fold anisotropy (12) with $A_1 = -0.015$,
 $p_0(x_1, x_2) = \sin(2x_1)\sin(2x_2)$ at different times

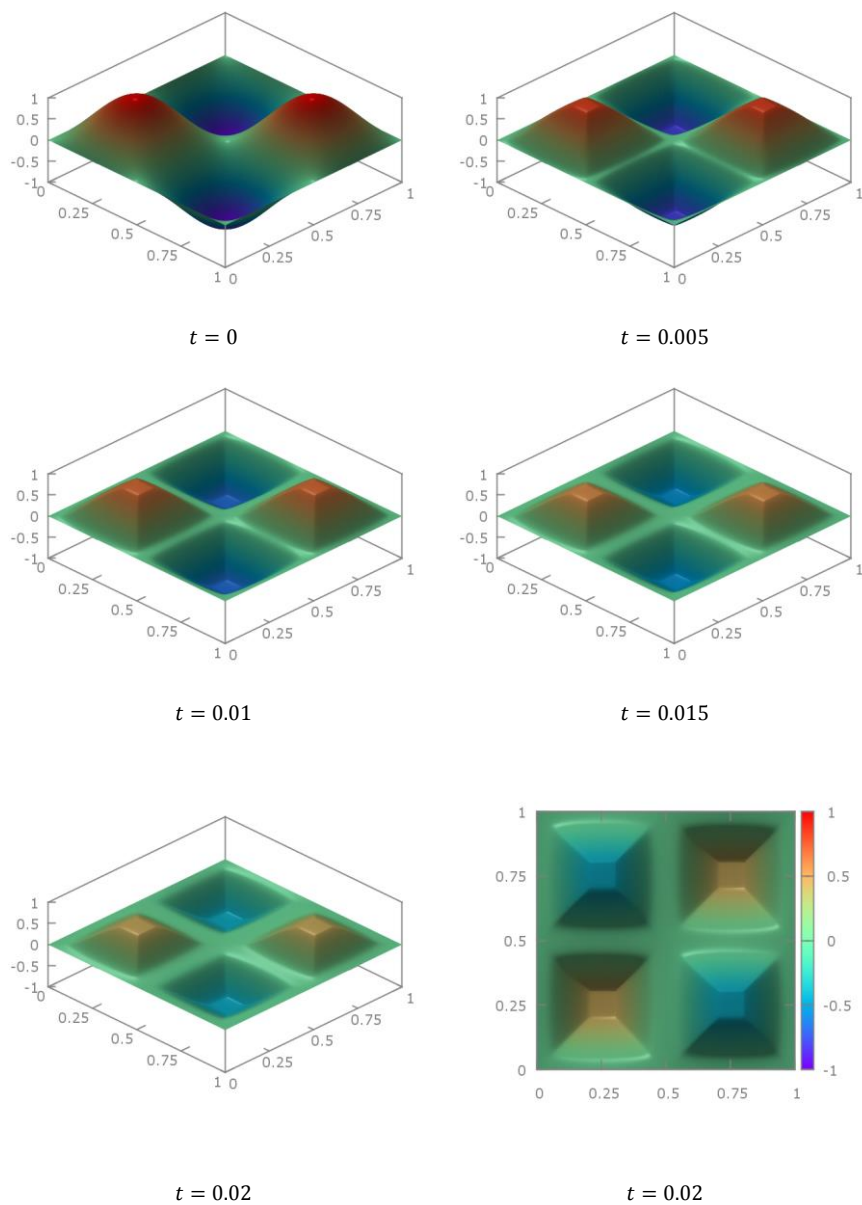


Figure 5

Morphology evolution for $f = 0$, the regularized l_1 norm (13) with $A_1 = 0.01$,
 $p_0(x_1, x_2) = \sin(2x_1)\sin(2x_2)$ at different times

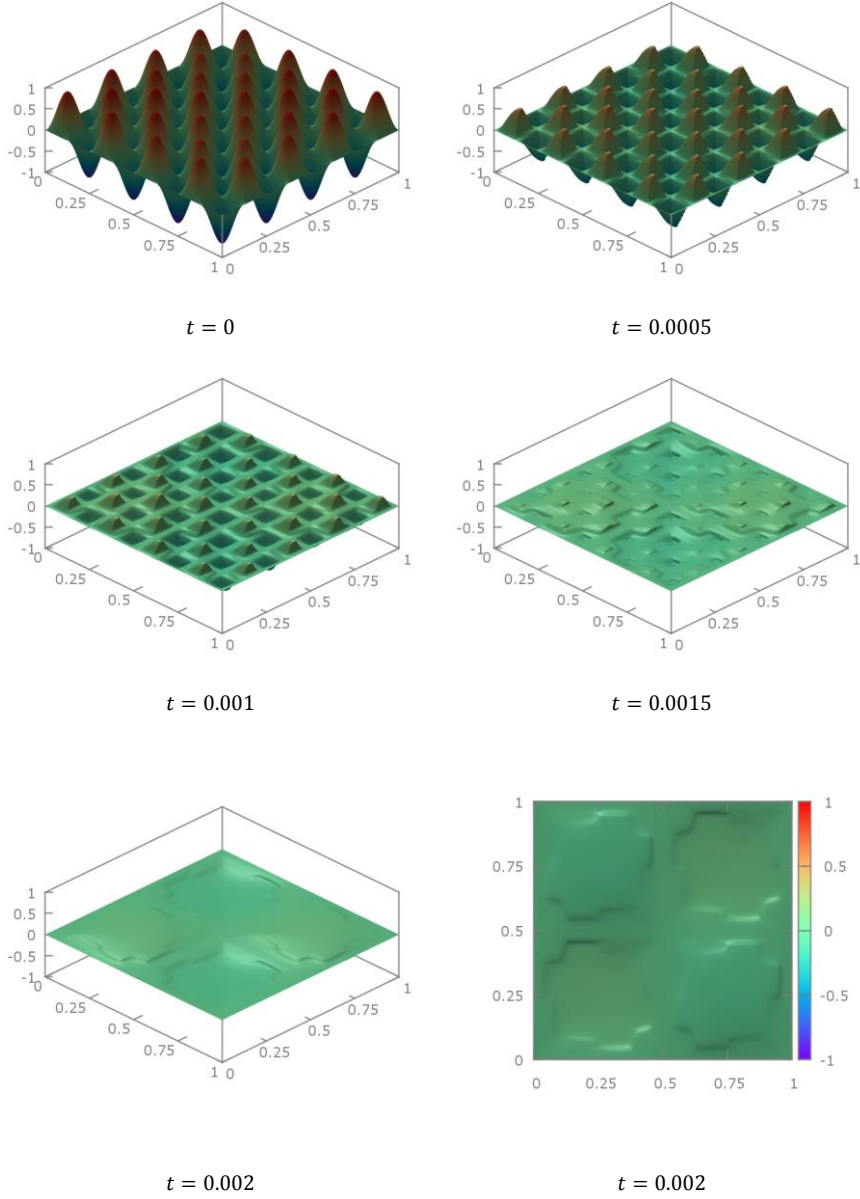


Figure 6

Morphology evolution for $f = 0$, the regularized l_1 norm (13) with $A_1 = 0.01$,
 $p_0(x_1, x_2) = 0.1\sin(2x_1)\sin(2x_2) + 0.9\sin(4x_1)\sin(4x_2)$ at different times

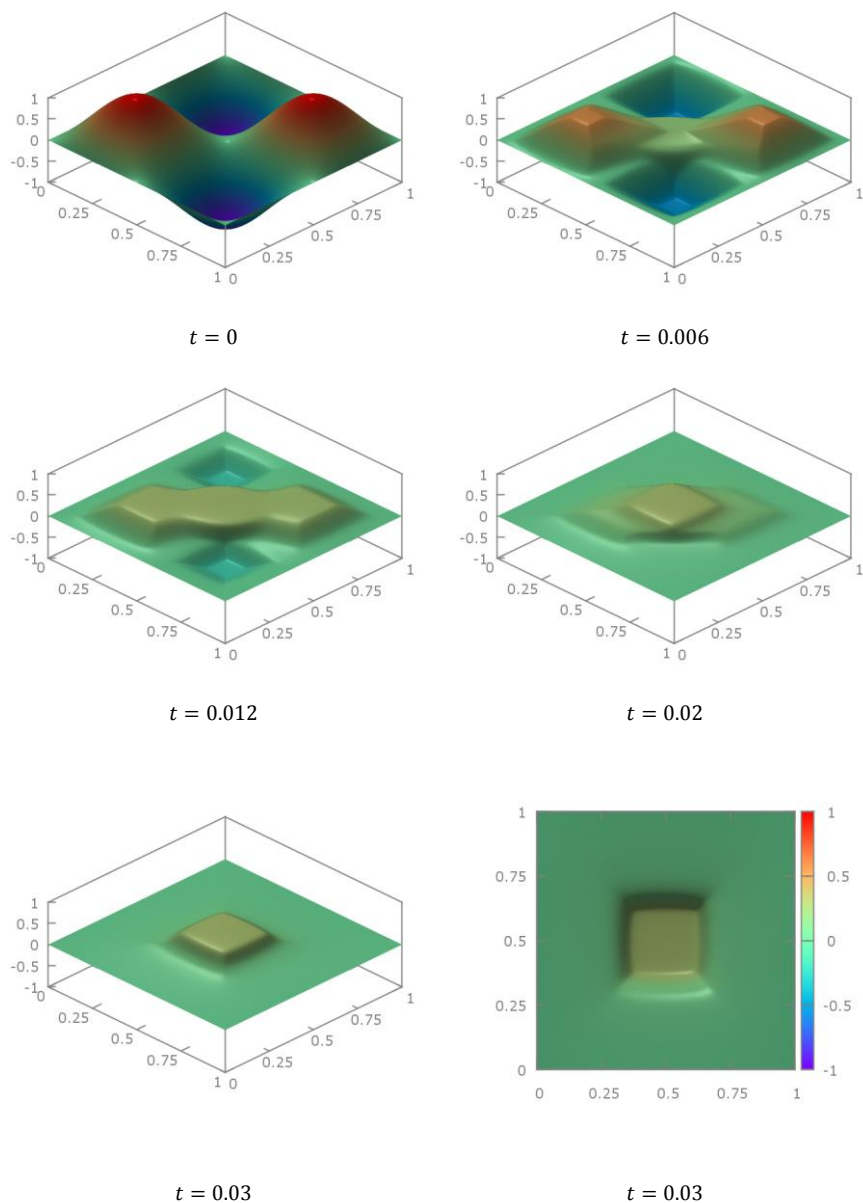


Figure 7

Morphology evolution for

$$f = 10(1 - \tanh(20(\sqrt{(x_1 - 0.5)(x_1 - 0.5) + (x_2 - 0.5)(x_2 - 0.5)} - 0.15))),$$

the regularized l_1 norm (13) with $A_1 = 0.01$, $p_0(x_1, x_2) = \sin(2x_1)\sin(2x_2)$ at different times

Conclusion

In the paper, we have studied the anisotropic mean curvature flow in relative geometry for which a global existence result has been derived. A numerical scheme based on the method of lines has been presented and analysed concerning its stability. In the numerical experiments, the influence of various anisotropy symmetries and the forcing term on the surface evolution has been addressed.

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